Quantitative Molecular Spectroscopy in Cavity

Presented by:

OSA Environmental Sensing Technical Group







Technical Group Leadership:

Krishnan Parameswaran, Analog Devices Inc., USA

Joachim Sacher, Sacher Lasertechnik GmbH, Germany

Amartya Sengupta, Indian Institute of Technology Delhi, India









Technical Group Website:

www.osa.org/EnvironmentalSensingTG



Over 1,100 Total Members

Scope:

This technical group covers optical tools and techniques used in environmental sensing, including DIAL and LIDAR, hyperspectral monitoring, detection, processing and characterization, surveying applications, atmospheric propagation, pollution monitoring, and remote imaging. Also included in this area is remote sensing for military and commercial applications such as land management, target detection, and disaster monitoring.



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Krishnan Parameswaran · Moderator 3d Photonics Engineering Manager at Analog Devices 3d Webinar! Volcanoes! 3d	Complement the OSA Technical Group
Hello All. I am happy to announce the first OSA Environmental Sensors Webinar of 2018. The exciting topic is Remote Sensing of Volcanoes using Smart Sensing Technology.	Member List
Here are the details Presenter: Dr. Andrew McGonigle, University of Sheffield Date: Wednesday January 17, 2018 Time: 1000 Eastern Time (United States), 1500 GMT	Activities: Webinars, Special Sessions in CLEO/Fi0
Registration Link: https://cc.callinfo.com/registration/#/?	

Welcome to Today's Webinar!



OSA Sensing Technical Group



QUANTITATIVE MOLECULAR SPECTROSCOPY IN CAVITY WEBINAR

6 December 2018 • 10:00 EST

Dr. Patrick Dupré, Université du Littoral Côte d'Opale

Patrick Dupré is a recognized expert in molecular high resolution and quantitative laser Spectroscopy. His career has included work in physics and chemistry laboratories in France, the United States, the United Kingdom and Germany. His interests include experimental spectroscopy and modeling. He is presently involved in developing Noise-Immune Cavity-Enhanced Optical Heterodyne Molecular Spectrometry (NICE-OHMS) for metrology applications and for trace gas detection in the Mid-InfraRed. Spectroscopy with high finesse cavity is an ideal tool for saturated absorption, i.e. under sub-Doppler conditions.

Quantitative Molecular Spectroscopy in Cavity

Patrick DUPRÉ

Laboratoire de Physico-Chimie de l'Atmosphère, ULCO Dunkerque, France December 2018



Outline

Motivations

- Introduction to CEAS
- 3 Absorption in Cavity
- Saturated Absorption: Modeling and Simulations
- Solution NICE-OHMS: Principles and Implementation
- \bigcirc C₂H₂ in the NIR
- 7 HD in the NIR (Forbidden Transition)
- Bibliography

Conclusions

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Basic Idea: **Enhancing the Absorption Length**, i.e., the length of interaction between light and analyte. How?

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Large Metallic Mirrors

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Off-Axis ICOS with recycling mirror to recover the leaking input power (with the permission of J. Mandon)

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• *S* is available in the database like HITRAN (http://hitran.org/) P. DUPRÉ (pdupre@gmx.com) (LPCA/ULCO) Spectroscopy in Cavity Dec. 2018



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$$I_{out}(t) = \left| FT^{-1} \Big[\mathcal{T}_{cav}(\omega) \cdot \mathcal{E}_{in}(\omega) \Big] \right|^2$$

Simulation: Pulsed Source



Simulation: CW Source



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- NICE-OHMS benefits of both: CW acquisition, and full noise Immunity.

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 - Simultaneous determination of the number density and of the crossection, from a single decay (CRDS)!

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NonExponential Decay

CRDS of Jet-Cooled NO₂: Decay of the line at 12536.4464 cm⁻¹ (${}^{q}R_{0}(0), 3/2$)



Saturation in NO₂ (with Fine and Hyperfine Transitions)

Power dependence of the ${}^{q_0}Q_{21}(0.5)$ Line Pattern



Saturation in NO₂ (with Fine and Hyperfine Transitions)

Power dependence of the ${}^{q_0}Q_{21}(0.5)$ Line Pattern Linear^a 0.6 W^a — Voigt 3e-06 1.5 W^a 0.5 W 1.5 W 3 Wb 3 W 3.7 W^a 6 W 10 W 20 W Absorption (/cm) 50 W 9.2 W^a 2e-06 100 W 250 W 19 W^b 46 W^b 1e-06 2 W^c Wc 12536.32 12536.305 12536.31 12536.315 Excitation Energy (cm⁻¹)

Power dependence of the ${}^{q_0}R_{11}(0.5)$ Line Pattern



Absorption versus Intracavity Power

Absorption versus the Intracavity Power at the Center of the ${}^{q}R_{11}(0.5)$ Line Pattern



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- Since 2017: Stefan Schäffer, Niels Bohr Institute (Copenhagen), **MOT** of ⁸⁸Sr (locking against transition)

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- 5 NICE-OHMS: Principles and Implementation
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- 7 HD in the NIR (Forbidden Transition)
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• Establishing the complex absorption in the Frequency domain (coupling of the EMF with the susceptibility) for a given Doppler Shift

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- Integration over the Transit-Time Rate (Maxwell Boltzmann)

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- Summation of the Degenerated Zeeman Sub-Transitions (Polarization)
- No Saturation Coefficient is used

Saturation Analysis: C₂H₂ Transition R(0) at 6558.79233 cm⁻¹ (polyad 10)



C₂H₂ NICE-OHMS Simulation (Absorption)

NICE-OHMS in Phase, C₂H₂ Transition R(0) at 7143.8289 cm⁻¹ (polyad 11)



C₂H₂ NICE-OHMS Simulation (Dispersion)



C₂H₂ Simulation (Dispersion): Lorentzian Component

Saturation Analysis: C₂H₂ Transition R(0) at 7143.8289 cm⁻¹ (polyad 11)



C₂H₂ Simulation (Dispersion): Gaussian Component

Saturation Analysis: C₂H₂ Transition R(0) at 7143.8289 cm⁻¹ (polyad 11)



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Spectroscopy in Cavity



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Immersed Cavity















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Direct Absorption of C₂H₂ (P 11, $v_1 + v_2 + (2v_4 + v_5)^1 \leftarrow 0$)



 L_{eq} : ~27.2 km

Direct Absorption of C₂H₂ (P 11, $v_1 + v_2 + (2v_4 + v_5)^1 \leftarrow 0$)



Direct Absorption of C₂H₂ (P 11, $v_1 + v_2 + (2v_4 + v_5)^1 \leftarrow 0$)



NICE-OHMS "Absorption" of C₂H₂ (Polyad 11)

 C_2H_2 , Transition $R_e(4)$ (7239.79077 cm⁻¹), NICE-OHMS in Phase at 7 µbar



NICE-OHMS "Dispersion" of C₂H₂ (Polyad 11)


NICE-OHMS Dispersion of C₂H₂ (Polyad 11) with OFC

C₂H₂, Transition R_e(4) (7239.79077 cm⁻¹), NICE-OHMS in Quadrature at 2 µbar



Resonance Width Power Dependence

Power Broadening, NICE-OHMS of C_2H_2 , Transition R_2 , $v_1 + v_2 + (2v_4 + v_5)^1$



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Simulation of HD: First Overtone, Transition R(0)



HD: 1st Overtone, Transition *R*(1)



HD Hyperfine Structure according to N. Ramsey



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 - Mean Transit-Time Rate: ~ 660 kHz

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- Improvement by 3 Orders of Magnitude
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 - Mean Transit-Time Rate: ~ 660 kHz
 - Rabi Frequency (~ 21 kHz)?

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- Resonance Line Shape (Asymmetry)?
 - Mean Transit-Time Rate: ~ 660 kHz
 - Rabi Frequency (~ 21 kHz)?
 - Recoil (~ 35 kHz)?

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- *R*(1), *R*(2), *R*(3) observed
- Sensitivity ~ 10^{-12} cm⁻¹
- Resonance Line Shape (Asymmetry)?
 - Mean Transit-Time Rate: ~ 660 kHz
 - Rabi Frequency (~ 21 kHz)?
 - Recoil (~ 35 kHz)?
 - Evidence of the Hyperfine Structure (spreads over 600 kHz)?

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Thank for your Attention





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Spectroscopy in Cavity