Quantum Measurement of Motion in the Negative Mass Reference Frame

Presented by:

OSA Quantum Computing and Communication Technical Group

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QUANTUM MEASUREMENT OF MOTION IN THE NEGATIVE MASS REFERENCE FRAME

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Quantum Computing and Communication Technical Group

Technical Group at a Glance

Focus

- •Theoretical and experimental aspects of quantum computing
- •Quantum communication systems Cryptography
- •Generation, detection and applications of non-classical light
- •Quantum measurement and quantum control

Mission

•To maximize the exchange of information and the creation of networking opportunities for our community

•Webinars, technical events (workshops, tutorials, poster sessions), outreach activities

•Interested in presenting your research? Have ideas for TG events? Contact us at <u>TGactivities@osa.org</u>.

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Quantum Measurement of Motion in the Negative Mass Reference Frame

Prof. Eugene Polzik

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Speaker's Short Bio:

Director of the Danish National Research Foundation Center for Quantum Optics (QUANTOP). Member of the Royal Danish Academy of Science. Fellow of the American Physical Society and The Optical Society. Recipient of several awards, including the Gordon Moore Distinguished Scholar award, the Scientific American Research Leadership Award and the Danish Association of Academics award.

Quantum measurement of motion in the <u>negative mass reference frame</u>

Eugene Polzik Niels Bohr Institute Copenhagen

Image credit Bastian Leonhardt Strube and Mads Vadsholt

European Research Council Established by the European Commission

Quantum limits for an oscillator

$Var(X) Var(P) \ge 1/4$

$X_{Lab}(t) = X \sin(\omega t) + P \cos(\omega t)$

The more you measure – the more you disturb!

The optimal measurement strength/duration allows to reach the Standard Quantum Limit (SQL) for position measurement:

$$Var(X_{lab}) = \frac{\hbar}{m\Omega}$$

And yet, arbitrary small perturbations in BOTH position and momentum can be measured simultaneously

Trajectories without quantum uncertainties

in negative mass reference frame

See also: Tsai and Caves, PRL 2010 M. Ozawa E.S. Polzik, K. Hammerer. Ann. der Physik 527, A15 (2015).
W. Wasilewski et al. PRL, 104, 133601 (2010).
K. Hammerer et al. PRL102, 020501 (2009).

Einstein-Podolsky-Rosen (EPR) entanglement 1935 $Var(X) Var(P) \ge 1/4$ 2 particles entangled $|\hat{X}_1 - \hat{X}_2, \hat{P}_1 + \hat{P}_2| = 0$ in position/momentum L \hat{X}_1, \hat{P}_1 \hat{X}_2, \hat{P}_2 $\hat{X}_1 - \hat{X}_2 = L$ $\hat{P}_1 + \hat{P}_2 = 0$

> Simon (2000); Duan, Giedke, Cirac, Zoller (2000) Necessary and sufficient condition for entanglement

$$Var(X - X_0) + Var(P + P_0) < 2$$

3 steps to noiseless quantum trajectories

1. Define trajectory relative to a quantum reference

2. Reference system has an effective negative mass

з. Entangled state of the reference and the probed systems ís generated

"Experimental long-lived entanglement of two macroscopic objects". B. Julsgaard, A. Kozhekin and ESP. **Nature**, 413, 400 (**2001**)

"Establishing Einstein-Podolsky-Rosen channels between nanomechanics and atomic ensembles". K. Hammerer, M. Aspelmeyer, ESP, P. Zoller. **PRL** 102, 020501 (**2009**).

"Trajectories without quantum uncertainties". K. Hammerer and ESP, Annalen der Physik . (2015)

Probe system entangled with origin system $X(dt)_{X0} = X(0)_{X0} + (\dot{X} - \dot{X}_0)dt$ $= X(0)_{X0} + (P \oplus P_0) dt$

Not good enough

 $m = \Theta m_0 = 1$

Oscíllators: negatíve mass=negatíve frequency

$$X(t) = X(0)\cos(\omega t) + P(0)\sin(\omega t)/m$$

For a reference oscillator with a negative frequency:

$$X(t) - X_0(t) = [X(0) - X_0(0)] \cos(\omega t) + [P(0) + P_0(0)] \sin(\omega t)$$

Q: how to implement a negative mass oscillator?

spin in magnetic field

Spin ensemble = negative mass harmonic oscillator

 $H = \chi_{Par} \hat{a}^{\dagger} \hat{b}^{\dagger} + \chi_{BS} \hat{a}^{\dagger} \hat{b} + h.c. = g X_S X_L,$ photon Polariton=collective spin

Quantum back-action-evading measurement of motion in a negative mass reference frame

Christoffer B. Møller¹*, Rodrigo A. Thomas¹*, Georgios Vasilakis^{1,2}, Emil Zeuthen^{1,3}, Yeghishe Tsaturyan¹, Mikhail Balabas^{1,4}, Kasper Jensen¹, Albert Schliesser¹, Klemens Hammerer³ & Eugene S. Polzik¹

LETTER doi:10.1038/nature22980

Treutlein group: Coupling atomic and mirror motion. PRL 2018

Distributed HYBRID quantum system of SPIN and MECHANICS

Michael Balabas

Room temperature spin quantum oscillator

Mechanical oscillator with $Q = 10^7 - 10^9$

Image credit Bastian Leonhardt Strube and Mads Vadsholt

C. B. Møller et al. LETTER doi:10.1038/nature22980

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Observation of Gravitational Waves from a Binary Black Hole Merger

B. P. Abbott *et al.** (LIGO Scientific Collaboration and Virgo Collaboration) (Received 21 January 2016; published 11 February 2016)

GWD BEYOND SQL WITH NEGATIVE MASS SPINS

2015-2017

- 2 observing runs: ∼ 1 year
- 5 BH+BH events
- 1 NS+NS event.

Advanced LIGO design goal

- ×2.5÷3 in sensitivity
- $\sim 10^2$ BH+BH events/year
- ≥ 10 NS+NS events/year

Still not enough for:

- Supernovae
- Pulsars
- Cosmic strings
- Background radiation

GWD end mirrors

suspended 40 kg "free" masses: oscillation frequency < 1Hz with Q = 10⁸

Standard Quantum Limit for measurement of motion of a free mass

$$X(t) = X + \frac{Pt}{m}, \quad \Delta X \Delta P \ge \frac{\hbar}{2} \Rightarrow$$
$$[\Delta X(t)]^2 \ge (\Delta X)^2 + \frac{\hbar^2 t^2}{4m^2 (\Delta X)^2} \ge \frac{\hbar t}{m} \quad (SQL)$$

Other proposals for beyond SQL Gravitational Wave Interferometry: H.J. Kimble et al, PRD65, 022002 (2001). Squeezed light with phase rotated by GWI Y.Ma et al, Nature Physics 13, 776 (2017). Entangled light injected into GWI

GWD BEYOND SQL WITH NEGATIVE MASS SPINS

F. Khalili and E.S.P. *Quantum back action evading detection of gravitational waves in a negative mass reference frame.* Phys. Rev. Lett. July 2018

Probing the hybrid system with EPR entangled light modes

$$\begin{split} P_{L1,out} &= -P_{L1,in} + force \ terms + \Gamma_M \chi_M X_{L1,in} \\ P_{L2,out} &= -P_{L2,in} + force \ terms + \Gamma_M \chi_M X_{L2,in} \\ \hline \mathbf{D1} \qquad \qquad \mathbf{Spins} \qquad \mathbf{852nm} \qquad \mathbf{GWI} \qquad \mathbf{D2} \\ \hline \mathbf{Fntangled} \qquad \mathbf{I064nm} \qquad \mathbf{GWI} \qquad \mathbf{D2} \\ P_{L1,out} - P_{L2,out} &= -P_{L1,in} + P_{L2,in} + force \ terms \\ &+ \Gamma_M \chi_M X_{L1,in} - \Gamma_S \chi_S X_{L2,in} \\ \hline \Gamma_M \chi_M &= -\Gamma_M \chi_M \qquad \qquad (P_{L1,in} - P_{L2,in})^2 = e^{-2r} \\ &\qquad (X_{L1,in} + X_{L2,in})^2 = e^{-2r} \\ P_{L1,out} - P_{L2,out} \Rightarrow force \ terms \end{split}$$

F. Khalili and E.S.P. Phys.Rev.Lett. July 2018

Simulation for LIGO

European Gravitational Wave future

Summary: standard quantum limits of measurement precision of fields and forces can be surpassed

Next generation sensors of force, acceleration, and gravity

Image credit: Bastian Leonhardt Strube and Mads Vadsholt